## Research Article

# New Examples of Einstein Metrics in Dimension Four 

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We construct new examples of four-dimensional Einstein metrics with neutral signature and twodimensional holonomy Lie algebra.

## 1. Introduction

The holonomy group of a metric $g$ at a point $p$ of a manifold $M$ is the group of all linear transformations in the tangent space of $p$ defined by parallel translation along all possible loops starting at $p$ [1]. It is obvious that a connection can only be the Levi-Civita connection of a metric $g$ if the holonomy group is a subgroup of the generalized orthogonal group corresponding to the signature of $g[1-3]$. At any point $p \in M$, and in some coordinate system about $p$, the set of matrices of the form

$$
\begin{equation*}
R_{j k l}^{i} X^{k} Y^{l}, R_{j k l ; m}^{i} X^{k} Y^{l} Z^{m}, R_{j k l ; m n}^{i} X^{k} Y^{l} Z^{m} W^{n}, \ldots, \tag{1.1}
\end{equation*}
$$

where $X, Y, Z, W \in T_{p} M$ and semicolon denotes covariant derivative, forms a Lie subalgebra of the Lie algebra of $M_{n}(\mathbb{R})$ of $\operatorname{GL}(n, \mathbb{R})$ called the infinitesimal holonomy algebra of $M$ at $p$. Up to isomorphism the latter is independent of the coordinate system chosen. The corresponding uniquely determined connected subgroup of $G L(n, \mathbb{R})$ is called the infinitesimal holonomy group of $M$ at $p$.

A metric tensor $g$ is a nondegenerate symmetric bilinear form on each tangent space $T_{p} M$ for all $p \in M$. The signature of a metric $g$ is the number of positive and negative eigenvalues of the metric $g$. The signature is denoted by an ordered pair of positive integers
$(p, q)$, where $p$ is the number of positive eigenvalues and $q$ is the number of negative eigenvalues. If $p=q$, we say that the metric is of neutral signature. In this article, we are interested in four-dimensional metrics with neutral signature.

If the metric $g$ satisfies the condition

$$
\begin{equation*}
R_{i j}=\frac{R}{4} g_{i j} \tag{1.2}
\end{equation*}
$$

where $R_{i j}$ are the components of the Ricci tensor and $R$ is the scalar curvature, then we say that $g$ is an Einstein metric and the pair $(M, g)$ is an Einstein space.

In [4], Ghanam and Thompson studied and classified the holonomy Lie subalgebras of neutral metrics in dimension four. In this paper, we will focus on one of the subalgebras presented in [4], namely, $A_{17}$. For this subalgebra we will show that the metric presented in [4] will lead us to the construction of Einstein metrics. In Section 3, we will give the metrics explicitly, and in Section 4, we will show that these Einstein metrics produce $A_{17}$ at their two-dimensional holonomy.

As a final remark regarding our notation, we will use subscripts for partial derivatives. For example, the partial derivative of a function $a$ with respect to $x$ will be denoted by $a_{x}$.

## 2. The Subalgebra $A_{17}$ As a Holonomy

In this section we will consider the Lie algebra $A_{17}$; it is a 2-dimensional Lie subalgebra of the Lie algebra of the generalized orthogonal group $O(2,2)$ of neutral signature [4,5]. A basis for $A_{17}$ is given by

$$
e_{1}=\left[\begin{array}{cc}
-J & J  \tag{2.1}\\
-J & J
\end{array}\right], \quad e_{2}=\left[\begin{array}{cc}
J & L \\
L & J
\end{array}\right],
$$

where

$$
J=\left[\begin{array}{cc}
0 & 1  \tag{2.2}\\
-1 & 0
\end{array}\right], \quad L=\left[\begin{array}{ll}
0 & 1 \\
1 & 0
\end{array}\right]
$$

We turn now to a theorem of Walker [6] that will be a key to the existence of a metric $g$ that produces $A_{17}$ as a two-dimensional holonomy.

Theorem 2.1 (Walker [6]). Let $(M, g)$ be a pseudo-Riemannian manifold of class $C^{\infty}$. If $g$ admits a parallel, null $r$-distribution, then there is a system of coordinates $\left(x^{i}\right)$ relative to which $g$ assumes the following form:

$$
g_{i j}=\left[\begin{array}{ccc}
0 & 0 & I  \tag{2.3}\\
0 & A & H \\
I & H^{t} & B
\end{array}\right],
$$

where $I$ is the $r \times r$ identity matrix and $A, B, H$, and $H^{t}$ are matrix functions of the same class as $M$, satisfying the following conditions but otherwise arbitrary.
(1) $A$ and $B$ are symmetric; $A$ is of order $(n-2 r) \times(n-2 r)$ and nonsingular, $B$ is of order $r \times r, H$ is of order $(n-2 r) \times r$, and $H^{t}$ is the transpose of $H$.
(2) $A$ and $H$ are independent of the coordinates $x^{1}, x^{2}, \ldots, x^{r}$.

Now we show that $A_{17}$ is a holonomy Lie algebra of a four-dimensional neutral metric.
Proposition 2.2. $A_{17}$ is a holonomy algebra.
Proof. In this case, we have an invariant null 2-distribution, and so by Walker's theorem, there exists a coordinate system, say $(x, y, z, w)$, such that the metric $g$ is of the form

$$
g=\left[\begin{array}{llll}
0 & 0 & 1 & 0  \tag{2.4}\\
0 & 0 & 0 & 1 \\
1 & 0 & a & c \\
0 & 1 & c & b
\end{array}\right]
$$

where $a, b, c$ are smooth functions in $(x, y, z, w)$. Since the invariant distribution contains a parallel null vector field, we must have

$$
\begin{equation*}
a_{x}=b_{x}=c_{x}=0 \tag{2.5}
\end{equation*}
$$

It was shown in [4] that, in order for $g$ to produce $A_{17}$ as its holonomy algebra, the functions $a, b$, and $c$ must satisfy the following conditions:

$$
\begin{equation*}
b_{y y}=0, \quad c_{y y}=0, \quad c_{y w}-b_{y z}=0 \tag{2.6}
\end{equation*}
$$

## 3. New Einstein Metrics

In Section 2, we obtained a metric of the form

$$
g=\left[\begin{array}{llll}
0 & 0 & 1 & 0  \tag{3.1}\\
0 & 0 & 0 & 1 \\
1 & 0 & a & c \\
0 & 1 & c & b
\end{array}\right]
$$

where $a, b$, and $c$ are smooth functions in $y, z, w$ and they satisfy (2.5) and (2.6). We solve these conditions to obtain

$$
\begin{gather*}
b(y, z, w)=m(z, w) y+n(z, w), \\
c(y, z, w)=r(z, w) y+s(z, w) \tag{3.2}
\end{gather*}
$$

with

$$
\begin{equation*}
r_{w}=m_{z} \tag{3.3}
\end{equation*}
$$

where $m, n, r, s$ are smooth functions in $z, w$.
The nonzero components of the Ricci tensor for $g$ are

$$
\begin{gather*}
R_{33}=-\frac{y}{2} m a_{y y}-\frac{1}{2} n a_{y y}+a_{y w}-r_{z}-\frac{1}{2} m a_{y}+\frac{1}{2} r^{2} \\
R_{34}=\frac{1}{2}\left(r_{w}-m_{z}\right) \tag{3.4}
\end{gather*}
$$

The Ricci scalar is

$$
\begin{equation*}
R=0 \tag{3.5}
\end{equation*}
$$

Because of (3.3), we obtain

$$
\begin{equation*}
R_{34}=0 \tag{3.6}
\end{equation*}
$$

In this case, the Einstein condition $R_{i j}=(R / 4) g_{i j}$ becomes

$$
\begin{equation*}
R_{i j}=0 \tag{3.7}
\end{equation*}
$$

Hence, in order to have an Einstein metric, we must have

$$
\begin{equation*}
R_{33}=0 \tag{3.8}
\end{equation*}
$$

and so we obtain the following partial differential equation (PDE):

$$
\begin{equation*}
-\frac{y}{2} m a_{y y}-\frac{1}{2} n a_{y y}+a_{y w}-r_{z}-\frac{1}{2} m a_{y}+\frac{1}{2} r^{2}=0 \tag{3.9}
\end{equation*}
$$

Since we are interested in finding at least one solution, we take the following special values in (3.9):

$$
\begin{equation*}
m=1, \quad n=0, \quad r=0, \quad s=0 \tag{3.10}
\end{equation*}
$$

to obtain the following PDE:

$$
\begin{equation*}
y a_{y y}-2 a_{y w}+a_{y}=0 \tag{3.11}
\end{equation*}
$$

To solve (3.11), we use the method of separation. For example, assume that $a(y, z, w)$ is of the form

$$
\begin{equation*}
a(y, z, w)=f(y) g(z) h(w) \tag{3.12}
\end{equation*}
$$

where $f, g$, and $h$ are smooth functions in $y, z$, and $w$, respectively. We substitute (3.12) in (3.11) to obtain

$$
\begin{equation*}
y f^{\prime \prime}(y) g(z) h(w)-2 f^{\prime}(y) g(z) h^{\prime}(w)-f^{\prime}(y) g(z) h(w)=0 . \tag{3.13}
\end{equation*}
$$

We assume that $g(z)$ is nowhere zero to obtain

$$
\begin{equation*}
y f^{\prime \prime}(y) h(w)-2 f^{\prime}(y) h^{\prime}(w)-f^{\prime}(y) h(w)=0, \tag{3.14}
\end{equation*}
$$

and so

$$
\begin{equation*}
y f^{\prime \prime}(y) h(w)-f^{\prime}(y)\left(2 h^{\prime}(w)-h(w)\right)=0 . \tag{3.15}
\end{equation*}
$$

Hence

$$
\begin{equation*}
y f^{\prime \prime}(y) h(w)=f^{\prime}(y)\left(2 h^{\prime}(w)-h(w)\right) . \tag{3.16}
\end{equation*}
$$

Dividing both sides by $h(w) f^{\prime}(y)$ gives

$$
\begin{equation*}
\frac{y f^{\prime \prime}(y)}{f^{\prime}(y)}=\frac{2 h^{\prime}(w)-h(w)}{h(w)}=c, \tag{3.1.}
\end{equation*}
$$

where $c$ is a constant.
We now solve (3.17) and for that we will consider three cases.
(1) If $c=0$, then $f$ is a linear function given by

$$
\begin{equation*}
f(y)=c_{1} y+c_{2}, \tag{3.18}
\end{equation*}
$$

and the condition on $h$ becomes

$$
\begin{equation*}
2 h^{\prime}-h=0, \tag{3.19}
\end{equation*}
$$

which gives

$$
\begin{equation*}
h(w)=c_{3} e^{w / 2} . \tag{3.20}
\end{equation*}
$$

The solution $a(y, z, w)$ to the PDE equation (3.11) is

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} y+c_{2}\right) c_{4} e^{w / 2} g(z)=\left(c_{1} y+c_{2}\right) e^{w / 2} g(z) \tag{3.21}
\end{equation*}
$$

where $c_{1}, c_{2}$ are constants and $g(z)$ is a smooth nowhere zero function.
(2) If $c \neq 0,-1$, then the differential equations (3.17) become

$$
\begin{gather*}
y f^{\prime \prime}-c f^{\prime}=0  \tag{3.22}\\
2 h^{\prime}-(1+c) h=0
\end{gather*}
$$

The solutions to (3.22) are

$$
\begin{align*}
& f(y)=c_{1} y^{c+1}+c_{2} \\
& h(w)=c_{3} e^{(1+c) w / 2} \tag{3.23}
\end{align*}
$$

and so

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} y^{c+1}+c_{2}\right) e^{(1+c) w / 2} g(z) \tag{3.24}
\end{equation*}
$$

where $g(z)$ is a no-where zero smooth function in $z$.
(3) If $c=-1$, then the differential equations (3.17) become

$$
\begin{gather*}
y f^{\prime \prime}+f^{\prime}=0 \\
h^{\prime}=0 . \tag{3.25}
\end{gather*}
$$

The solutions to (3.25) are

$$
\begin{gather*}
f(y)=c_{1} \ln (y)+c_{2}  \tag{3.26}\\
h(w)=c_{3}
\end{gather*}
$$

and so

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} \ln (y)+c_{2}\right) c_{3} g(z)=\left(c_{1} \ln (y)+c_{2}\right) g(z) \tag{3.27}
\end{equation*}
$$

where $c_{1}, c_{2}$ are constants and $g(z)$ is a no-where zero smooth function in $z$.

## 4. The Holonomy of the New Metrics

In this section we compute the infinitesimal holonomy algebra and make sure that it produces a two-dimensional algebra. To do so, we consider our metric $g$ given by

$$
g=\left[\begin{array}{cccc}
0 & 0 & 1 & 0  \tag{4.1}\\
0 & 0 & 0 & 1 \\
1 & 0 & a(y, z, w) & 0 \\
0 & 1 & 0 & y
\end{array}\right]
$$

The only nonzero components of the curvature are

$$
\begin{gather*}
R_{2323}=-\frac{1}{2} a_{y y}, \\
R_{2334}=\frac{1}{2} a_{w y}-\frac{1}{4} a_{y},  \tag{4.2}\\
R_{3434}=-\frac{1}{2} a_{w w}+\frac{y}{4} a_{y}-\frac{1}{4} a_{w} .
\end{gather*}
$$

The holonomy matrices are

$$
\begin{gather*}
R_{i j 23}=\left[\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & -\frac{1}{2} a_{y y} & 0 \\
0 & -\frac{1}{2} a_{y y} & 0 & 0 \\
0 & 0 & 0 & 0
\end{array}\right], \\
R_{i j 34}\left[\begin{array}{cccc}
0 & 0 & 0 & 0 \\
0 & 0 & \frac{1}{2} a_{w y}-\frac{1}{4} a_{y} & 0 \\
0 & 0 & -\frac{1}{2} a_{w w}+\frac{y}{4} a_{y}-\frac{1}{4} a_{w} \\
0 & 0 & -\frac{1}{2} a_{w w}+\frac{y}{4} a_{y}-\frac{1}{4} a_{w} & 0
\end{array}\right], \tag{4.3}
\end{gather*}
$$

Now, in order for the metric to produce two-dimensional holonomy, we must have

$$
\begin{gather*}
a_{y y} \neq 0  \tag{4.4}\\
-\frac{1}{2} a_{w w}+\frac{y}{4} a_{y}-\frac{1}{4} a_{w} \neq 0 \tag{4.5}
\end{gather*}
$$

We have to check these equations for the three cases discussed in Section 3.
(1) We consider

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} y+c_{2}\right) e^{w / 2} g(z) \tag{4.6}
\end{equation*}
$$

Then

$$
\begin{equation*}
a_{y y}=0 \tag{4.7}
\end{equation*}
$$

and we obtain a one-dimensional holonomy. Therefore we must exclude this case.
(2) We consider

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} y^{c+1}+c_{2}\right) e^{(1+c) w / 2} g(z) \tag{4.8}
\end{equation*}
$$

where $c \neq 0,-1$. In this case

$$
\begin{equation*}
a_{y y}=c_{1} c(c+1) y^{c-1} e^{(1+c) w / 2} g(z) \neq 0 \tag{4.9}
\end{equation*}
$$

The second condition is that (4.5) gives

$$
\begin{equation*}
\frac{e^{(c+1) w / 2}}{8}\left(2 c_{1} y^{c+1} c^{2}+5 c_{1} y^{c+1} c+3 c_{1} y^{c+1}+2 c_{2} c^{2}+3 c_{2} c+c_{2}\right) \neq 0 \tag{4.10}
\end{equation*}
$$

This shows that the metric we constructed in Section 3 is an Einstein metric with a two-dimensional holonomy. In fact, its holonomy Lie algebra is $A_{17}$.
(3) We consider

$$
\begin{equation*}
a(y, z, w)=\left(c_{1} \ln (y)+c_{2}\right) g(z) . \tag{4.11}
\end{equation*}
$$

In this case the first condition is that (4.4) gives

$$
\begin{equation*}
a_{y y}=-\frac{c_{1} g(z)}{y^{2}} \neq 0 \tag{4.12}
\end{equation*}
$$

and the second condition is that (4.5) gives

$$
\begin{equation*}
\frac{c_{1} g(z)}{4} \neq 0 \tag{4.13}
\end{equation*}
$$

And once again we obtain an Einstein metric with $A_{17}$ as its two-dimensional holonomy Lie algebra.

## References

[1] G. S. Hall, "Connections and symmetries in spacetime," General Relativity and Gravitation, vol. 20, no. 4, pp. 399-406, 1988.
[2] G. S. Hall and W. Kay, "Holonomy groups in general relativity," Journal of Mathematical Physics, vol. 29, no. 2, pp. 428-432, 1988.
[3] G. S. Hall and W. Kay, "Curvature structure in general relativity," Journal of Mathematical Physics, vol. 29, no. 2, pp. 420-427, 1988.
[4] R. Ghanam and G. Thompson, "The holonomy Lie algebras of neutral metrics in dimension four," Journal of Mathematical Physics, vol. 42, no. 5, pp. 2266-2284, 2001.
[5] R. Ghanam, "A note on the generalized neutral orthogonal group in dimension four," Tamkang Journal of Mathematics, vol. 37, no. 1, pp. 93-103, 2006.
[6] A. G. Walker, "Canonical form for a Riemannian space with a parallel field of null planes," The Quarterly Journal of Mathematics. Oxford Second Series, vol. 1, pp. 69-79, 1950.

