Hindawi Publishing Corporation International Journal of Mathematics and Mathematical Sciences Volume 2010, Article ID 716035, 9 pages doi:10.1155/2010/716035

Research Article

New Examples of Einstein Metrics in Dimension Four

Ryad Ghanam

Department of Mathematics, University of Pittsburgh at Greensburg, 150 Finoli Drive, Greensburg, PA 15601, USA

Correspondence should be addressed to Ryad Ghanam, ghanam@pitt.edu

Received 27 January 2010; Accepted 23 March 2010

Academic Editor: Daniele C. Struppa

Copyright © 2010 Ryad Ghanam. This is an open access article distributed under the Creative Commons Attribution License, which permits unrestricted use, distribution, and reproduction in any medium, provided the original work is properly cited.

We construct new examples of four-dimensional Einstein metrics with neutral signature and twodimensional holonomy Lie algebra.

1. Introduction

The holonomy group of a metric g at a point p of a manifold M is the group of all linear transformations in the tangent space of p defined by parallel translation along all possible loops starting at p [1]. It is obvious that a connection can only be the Levi-Civita connection of a metric g if the holonomy group is a subgroup of the generalized orthogonal group corresponding to the signature of g [1–3]. At any point $p \in M$, and in some coordinate system about p, the set of matrices of the form

$$R^{i}_{jkl}X^{k}Y^{l}, R^{i}_{jkl;m}X^{k}Y^{l}Z^{m}, R^{i}_{jkl;mn}X^{k}Y^{l}Z^{m}W^{n}, \dots,$$

$$(1.1)$$

where $X, Y, Z, W \in T_p M$ and semicolon denotes covariant derivative, forms a Lie subalgebra of the Lie algebra of $M_n(\mathbb{R})$ of $GL(n, \mathbb{R})$ called the infinitesimal holonomy algebra of M at p. Up to isomorphism the latter is independent of the coordinate system chosen. The corresponding uniquely determined connected subgroup of $GL(n, \mathbb{R})$ is called the infinitesimal holonomy group of M at p.

A metric tensor g is a nondegenerate symmetric bilinear form on each tangent space T_pM for all $p \in M$. The signature of a metric g is the number of positive and negative eigenvalues of the metric g. The signature is denoted by an ordered pair of positive integers

(p,q), where *p* is the number of positive eigenvalues and *q* is the number of negative eigenvalues. If p = q, we say that the metric is of neutral signature. In this article, we are interested in four-dimensional metrics with neutral signature.

If the metric g satisfies the condition

$$R_{ij} = \frac{R}{4}g_{ij},\tag{1.2}$$

where R_{ij} are the components of the Ricci tensor and R is the scalar curvature, then we say that g is an Einstein metric and the pair (M, g) is an Einstein space.

In [4], Ghanam and Thompson studied and classified the holonomy Lie subalgebras of neutral metrics in dimension four. In this paper, we will focus on one of the subalgebras presented in [4], namely, A_{17} . For this subalgebra we will show that the metric presented in [4] will lead us to the construction of Einstein metrics. In Section 3, we will give the metrics explicitly, and in Section 4, we will show that these Einstein metrics produce A_{17} at their two-dimensional holonomy.

As a final remark regarding our notation, we will use subscripts for partial derivatives. For example, the partial derivative of a function *a* with respect to *x* will be denoted by a_x .

2. The Subalgebra A₁₇ As a Holonomy

In this section we will consider the Lie algebra A_{17} ; it is a 2-dimensional Lie subalgebra of the Lie algebra of the generalized orthogonal group O(2, 2) of neutral signature [4, 5]. A basis for A_{17} is given by

$$e_1 = \begin{bmatrix} -J & J \\ -J & J \end{bmatrix}, \qquad e_2 = \begin{bmatrix} J & L \\ L & J \end{bmatrix}, \tag{2.1}$$

where

$$J = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix}, \qquad L = \begin{bmatrix} 0 & 1 \\ 1 & 0 \end{bmatrix}.$$
 (2.2)

We turn now to a theorem of Walker [6] that will be a key to the existence of a metric g that produces A_{17} as a two-dimensional holonomy.

Theorem 2.1 (Walker [6]). Let (M, g) be a pseudo-Riemannian manifold of class C^{∞} . If g admits a parallel, null r-distribution, then there is a system of coordinates (x^i) relative to which g assumes the following form:

$$g_{ij} = \begin{bmatrix} 0 & 0 & I \\ 0 & A & H \\ I & H^t & B \end{bmatrix},$$
 (2.3)

where I is the $r \times r$ identity matrix and A, B, H, and H^t are matrix functions of the same class as M, satisfying the following conditions but otherwise arbitrary.

- (1) A and B are symmetric; A is of order $(n 2r) \times (n 2r)$ and nonsingular, B is of order $r \times r$, H is of order $(n 2r) \times r$, and H^t is the transpose of H.
- (2) A and H are independent of the coordinates x^1, x^2, \ldots, x^r .

Now we show that A_{17} is a holonomy Lie algebra of a four-dimensional neutral metric.

Proposition 2.2. A_{17} is a holonomy algebra.

Proof. In this case, we have an invariant null 2-distribution, and so by Walker's theorem, there exists a coordinate system, say (x, y, z, w), such that the metric g is of the form

$$g = \begin{bmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 1 & 0 & a & c \\ 0 & 1 & c & b \end{bmatrix},$$
 (2.4)

where a, b, c are smooth functions in (x, y, z, w). Since the invariant distribution contains a parallel null vector field, we must have

$$a_x = b_x = c_x = 0. (2.5)$$

It was shown in [4] that, in order for *g* to produce A_{17} as its holonomy algebra, the functions *a*, *b*, and *c* must satisfy the following conditions:

$$b_{yy} = 0, \qquad c_{yy} = 0, \qquad c_{yw} - b_{yz} = 0.$$
 (2.6)

3. New Einstein Metrics

In Section 2, we obtained a metric of the form

$$g = \begin{bmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 1 & 0 & a & c \\ 0 & 1 & c & b \end{bmatrix},$$
(3.1)

where a, b, and c are smooth functions in y, z, w and they satisfy (2.5) and (2.6). We solve these conditions to obtain

$$b(y, z, w) = m(z, w)y + n(z, w), c(y, z, w) = r(z, w)y + s(z, w)$$
(3.2)

with

$$r_w = m_z, \tag{3.3}$$

where *m*, *n*, *r*, *s* are smooth functions in *z*, *w*.

The nonzero components of the Ricci tensor for g are

$$R_{33} = -\frac{y}{2}ma_{yy} - \frac{1}{2}na_{yy} + a_{yw} - r_z - \frac{1}{2}ma_y + \frac{1}{2}r^2,$$

$$R_{34} = \frac{1}{2}(r_w - m_z).$$
(3.4)

The Ricci scalar is

$$R = 0. \tag{3.5}$$

Because of (3.3), we obtain

$$R_{34} = 0.$$
 (3.6)

In this case, the Einstein condition $R_{ij} = (R/4)g_{ij}$ becomes

$$R_{ij} = 0.$$
 (3.7)

Hence, in order to have an Einstein metric, we must have

$$R_{33} = 0,$$
 (3.8)

and so we obtain the following partial differential equation (PDE):

$$-\frac{y}{2}ma_{yy} - \frac{1}{2}na_{yy} + a_{yw} - r_z - \frac{1}{2}ma_y + \frac{1}{2}r^2 = 0.$$
(3.9)

Since we are interested in finding at least one solution, we take the following special values in (3.9):

$$m = 1, \quad n = 0, \quad r = 0, \quad s = 0$$
 (3.10)

to obtain the following PDE:

$$ya_{yy} - 2a_{yw} + a_y = 0. ag{3.11}$$

To solve (3.11), we use the method of separation. For example, assume that a(y, z, w) is of the form

$$a(y,z,w) = f(y)g(z)h(w), \qquad (3.12)$$

where f, g, and h are smooth functions in y, z, and w, respectively. We substitute (3.12) in (3.11) to obtain

$$yf''(y)g(z)h(w) - 2f'(y)g(z)h'(w) - f'(y)g(z)h(w) = 0.$$
(3.13)

We assume that g(z) is nowhere zero to obtain

$$yf''(y)h(w) - 2f'(y)h'(w) - f'(y)h(w) = 0, (3.14)$$

and so

$$yf''(y)h(w) - f'(y)(2h'(w) - h(w)) = 0.$$
(3.15)

Hence

$$yf''(y)h(w) = f'(y)(2h'(w) - h(w)).$$
(3.16)

Dividing both sides by h(w)f'(y) gives

$$\frac{yf''(y)}{f'(y)} = \frac{2h'(w) - h(w)}{h(w)} = c,$$
(3.17)

We now solve (3.17) and for that we will consider three cases.

(1) If c = 0, then f is a linear function given by

$$f(y) = c_1 y + c_2, (3.18)$$

and the condition on *h* becomes

$$2h' - h = 0, (3.19)$$

which gives

$$h(w) = c_3 e^{w/2}. (3.20)$$

The solution a(y, z, w) to the PDE equation (3.11) is

$$a(y,z,w) = (c_1y + c_2)c_4e^{w/2}g(z) = (c_1y + c_2)e^{w/2}g(z), \qquad (3.21)$$

where c_1, c_2 are constants and g(z) is a smooth nowhere zero function.

(2) If $c \neq 0, -1$, then the differential equations (3.17) become

$$yf'' - cf' = 0,$$

 $2h' - (1 + c)h = 0.$
(3.22)

The solutions to (3.22) are

$$f(y) = c_1 y^{c+1} + c_2,$$

$$h(w) = c_3 e^{(1+c)w/2},$$
(3.23)

and so

$$a(y,z,w) = (c_1 y^{c+1} + c_2) e^{(1+c)w/2} g(z), \qquad (3.24)$$

where g(z) is a no-where zero smooth function in z.

(3) If c = -1, then the differential equations (3.17) become

$$yf'' + f' = 0,$$

 $h' = 0.$
(3.25)

The solutions to (3.25) are

$$f(y) = c_1 \ln(y) + c_2,$$

 $h(w) = c_3,$
(3.26)

and so

$$a(y, z, w) = (c_1 \ln(y) + c_2)c_3g(z) = (c_1 \ln(y) + c_2)g(z), \qquad (3.27)$$

where c_1, c_2 are constants and g(z) is a no-where zero smooth function in z.

International Journal of Mathematics and Mathematical Sciences

4. The Holonomy of the New Metrics

In this section we compute the infinitesimal holonomy algebra and make sure that it produces a two-dimensional algebra. To do so, we consider our metric *g* given by

$$g = \begin{bmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ 1 & 0 & a(y, z, w) & 0 \\ 0 & 1 & 0 & y \end{bmatrix}.$$
 (4.1)

The only nonzero components of the curvature are

$$R_{2323} = -\frac{1}{2}a_{yy},$$

$$R_{2334} = \frac{1}{2}a_{wy} - \frac{1}{4}a_{y},$$

$$R_{3434} = -\frac{1}{2}a_{ww} + \frac{y}{4}a_{y} - \frac{1}{4}a_{w}.$$
(4.2)

The holonomy matrices are

$$R_{ij23} = \begin{bmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & -\frac{1}{2}a_{yy} & 0 \\ 0 & -\frac{1}{2}a_{yy} & 0 & 0 \\ 0 & 0 & 0 & 0 \end{bmatrix},$$

$$R_{ij34} \begin{bmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & \frac{1}{2}a_{wy} - \frac{1}{4}a_{y} & 0 \\ 0 & \frac{1}{2}a_{wy} - \frac{1}{4}a_{y} & 0 & -\frac{1}{2}a_{ww} + \frac{y}{4}a_{y} - \frac{1}{4}a_{w} \\ 0 & 0 & -\frac{1}{2}a_{ww} + \frac{y}{4}a_{y} - \frac{1}{4}a_{w} & 0 \end{bmatrix},$$

$$(4.3)$$

Now, in order for the metric to produce two-dimensional holonomy, we must have

$$a_{yy} \neq 0, \tag{4.4}$$

$$-\frac{1}{2}a_{ww} + \frac{y}{4}a_y - \frac{1}{4}a_w \neq 0.$$
(4.5)

We have to check these equations for the three cases discussed in Section 3.

(1) We consider

$$a(y, z, w) = (c_1 y + c_2) e^{w/2} g(z).$$
(4.6)

Then

$$a_{yy} = 0,$$
 (4.7)

and we obtain a one-dimensional holonomy. Therefore we must exclude this case.(2) We consider

$$a(y,z,w) = (c_1 y^{c+1} + c_2) e^{(1+c)w/2} g(z),$$
(4.8)

where $c \neq 0, -1$. In this case

$$a_{yy} = c_1 c(c+1) y^{c-1} e^{(1+c)w/2} g(z) \neq 0.$$
(4.9)

The second condition is that (4.5) gives

$$\frac{e^{(c+1)w/2}}{8} \left(2c_1 y^{c+1} c^2 + 5c_1 y^{c+1} c + 3c_1 y^{c+1} + 2c_2 c^2 + 3c_2 c + c_2 \right) \neq 0.$$
(4.10)

This shows that the metric we constructed in Section 3 is an Einstein metric with a two-dimensional holonomy. In fact, its holonomy Lie algebra is A_{17} .

(3) We consider

$$a(y, z, w) = (c_1 \ln(y) + c_2)g(z).$$
(4.11)

In this case the first condition is that (4.4) gives

$$a_{yy} = -\frac{c_1 g(z)}{y^2} \neq 0, \tag{4.12}$$

and the second condition is that (4.5) gives

$$\frac{c_1 g(z)}{4} \neq 0. \tag{4.13}$$

And once again we obtain an Einstein metric with A_{17} as its two-dimensional holonomy Lie algebra.

References

- G. S. Hall, "Connections and symmetries in spacetime," *General Relativity and Gravitation*, vol. 20, no. 4, pp. 399–406, 1988.
- [2] G. S. Hall and W. Kay, "Holonomy groups in general relativity," *Journal of Mathematical Physics*, vol. 29, no. 2, pp. 428–432, 1988.
- [3] G. S. Hall and W. Kay, "Curvature structure in general relativity," *Journal of Mathematical Physics*, vol. 29, no. 2, pp. 420–427, 1988.
- [4] R. Ghanam and G. Thompson, "The holonomy Lie algebras of neutral metrics in dimension four," *Journal of Mathematical Physics*, vol. 42, no. 5, pp. 2266–2284, 2001.
- [5] R. Ghanam, "A note on the generalized neutral orthogonal group in dimension four," *Tamkang Journal* of *Mathematics*, vol. 37, no. 1, pp. 93–103, 2006.
- [6] A. G. Walker, "Canonical form for a Riemannian space with a parallel field of null planes," *The Quarterly Journal of Mathematics. Oxford Second Series*, vol. 1, pp. 69–79, 1950.